Astrophysical and Instrumental Noise

Sources: Transits

Sagan Workshop – July 19 2016

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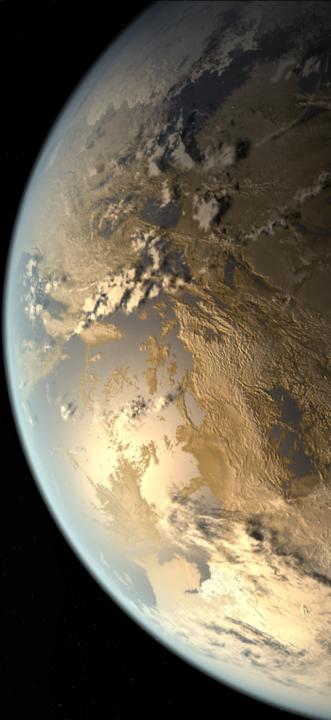
Why do we care about noise?

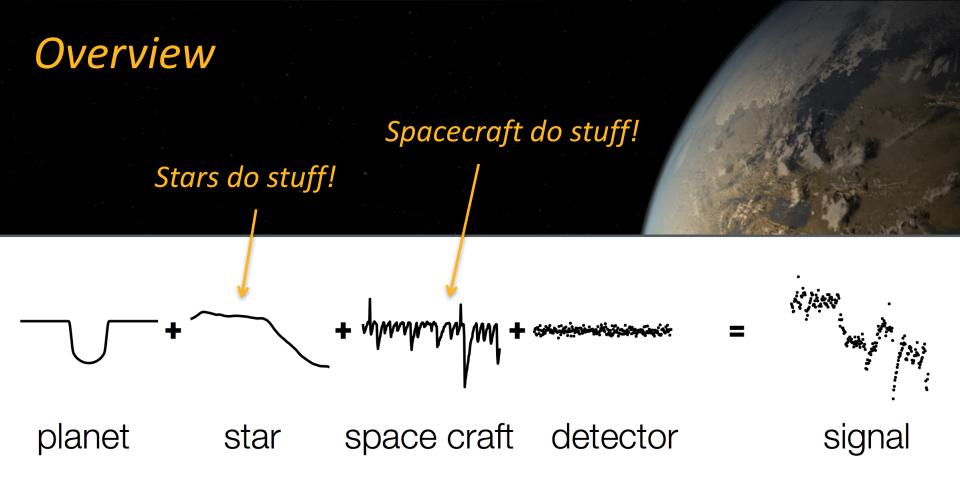
Take the example of the Kepler mission – the primary goal was determining the occurrence rate of Earth-like planets around sun-like stars

That's hard! Earth creates a 85ppm deep transit

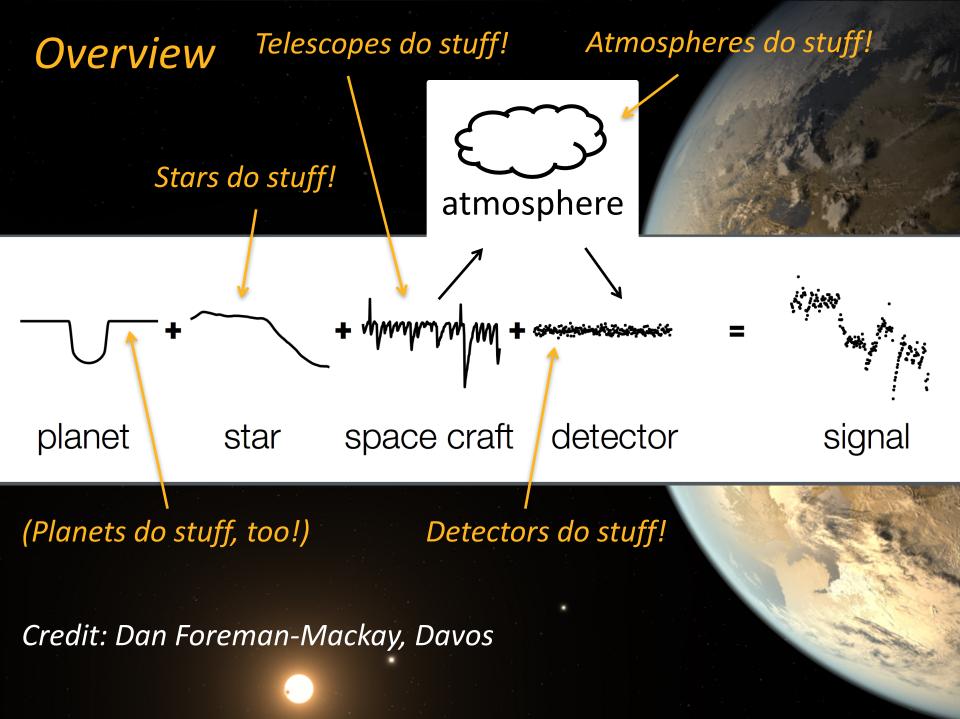
Mission design has to account for all the known sources of noise – error budget of 20ppm in 6.5h

How big a telescope do we need to build?
How faint a star can we look at?
What kind of stars can we look at?
How expensive a detector do we need to buy?
What is the overall noise floor achievable?





Credit: Dan Foreman-Mackay, Davos



Stars do stuff...

< 15 min

OSCILLATIONS OSCILLATIONS

a few m/s (Dumusque+ 11)

Kjeldsen+ 95, Bouchy & Carrier 01, Butler+ 04, Bedding & Kjeldsen 07

> MAGNETIC MAGNETIC CYCLES

1-20 m/s (Lovis+ 11)

Makarov 10, Dumusque+ 11 Dumusque+ 12, Meunier+ 13

~ 10 yrs

~ 1 h FLARES FLARES

<1 m/s (only active M)

Saar 09

STELLAR SIGNALS

Lindegren & Dravins 03

GRAVITATIONA GRAVITATIONAL REDSHIFT

< 10 cm/s (Cegla+12)

10 d - 10 yrs

15 min - 2 d

GRANULATION GRANULATION

a few m/s (Dumusque+ 11)

Del-Moro+ 04, Del-Moro 04 Cegla+ 12, Cegla+ 14

ACTIVE REGIONS

a few m/s (Meunier+ 10)

Saar & Donahue 97, Queloz+ 01 Hatzes 02, Meunier+ 10, Boisse+ 11, Dumusque+ 11, Lanza+ 11, Aigrain+12, Boisse+ 12, Reiners+ 13, Dumusque+ 14, Haywood+ 14,

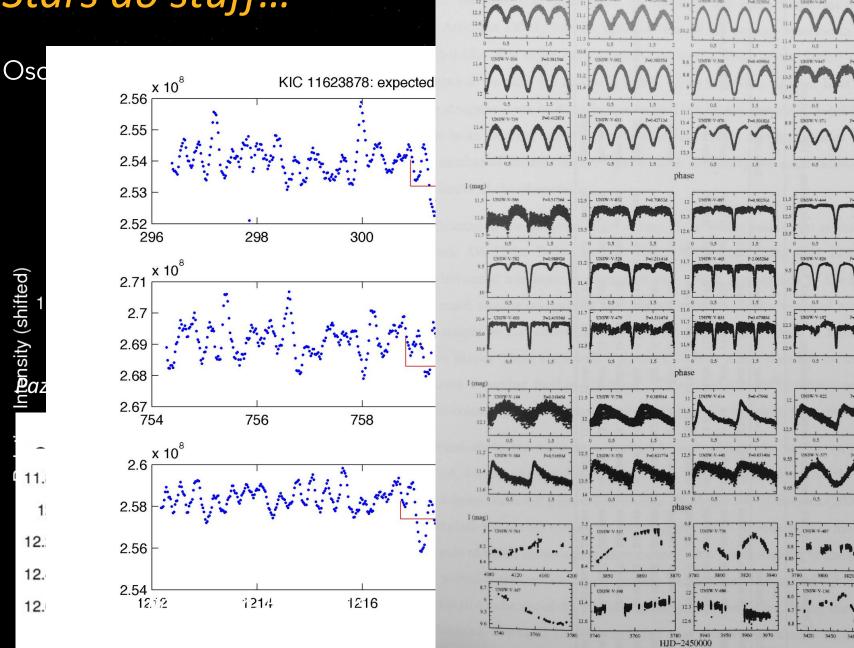
Rajpaul+ 15, Haywood+ 16

10 - 50 d

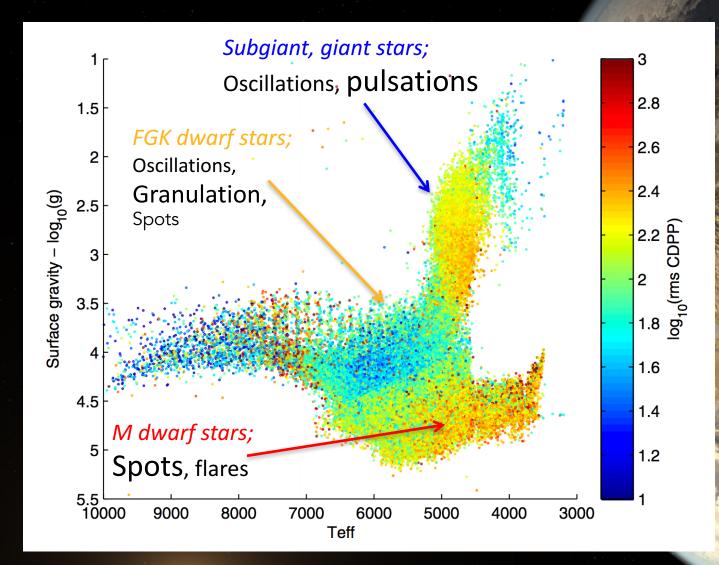
Credit: Xavier Dùmusque

Stars do stuff...



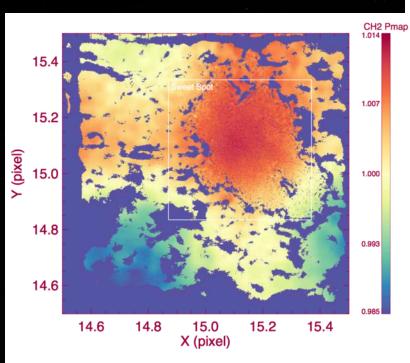


Different stars do different stuff...



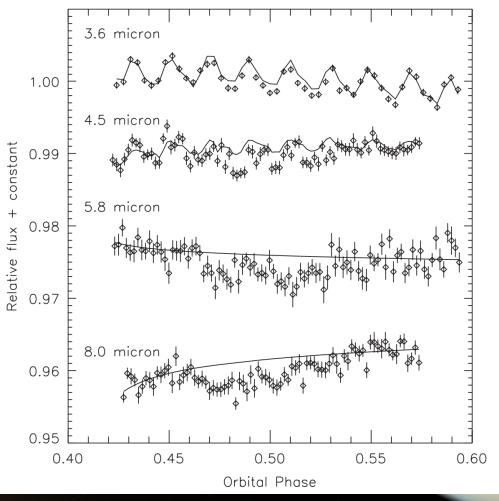
They typically have pointing jitter

• Intra-pixel variations, e.g. Spitzer, K2 (not Kepler!)



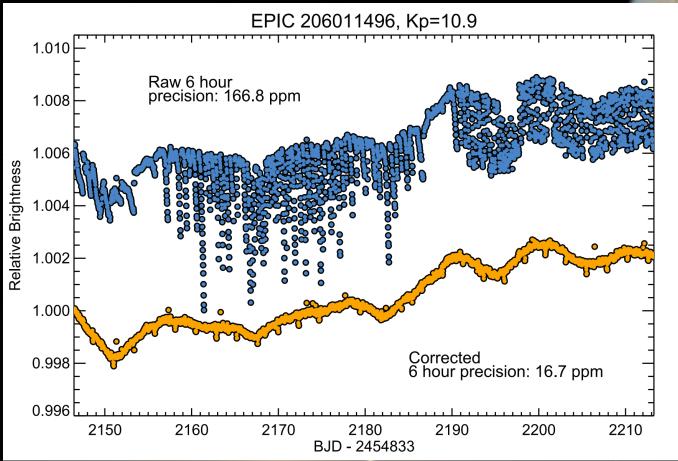
irachpp.spitzer.caltech.edu

Christiansen+2010



They typically have pointing jitter

• Intra-pixel variations, e.g. Spitzer, K2 (not Kepler!)

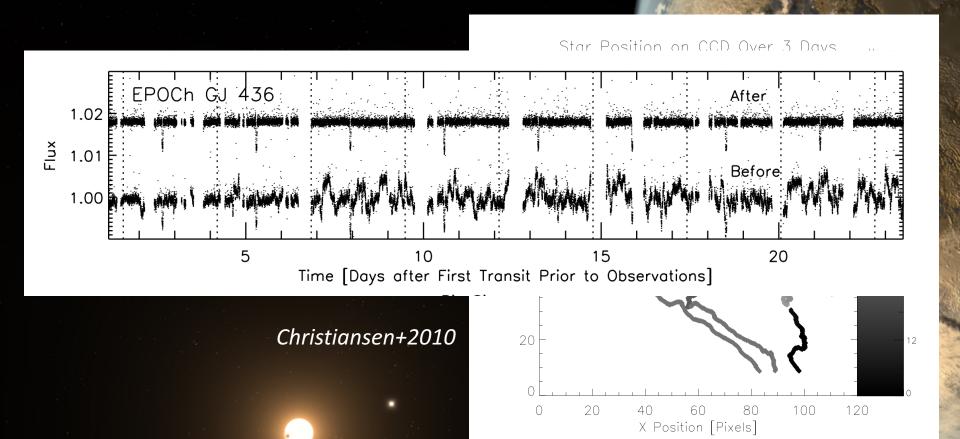


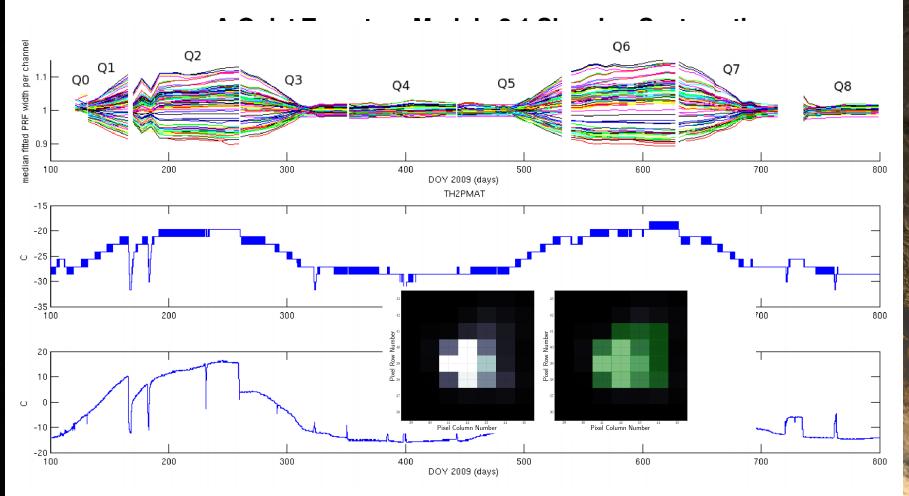


Vanderburg+2016

They typically have pointing jitter

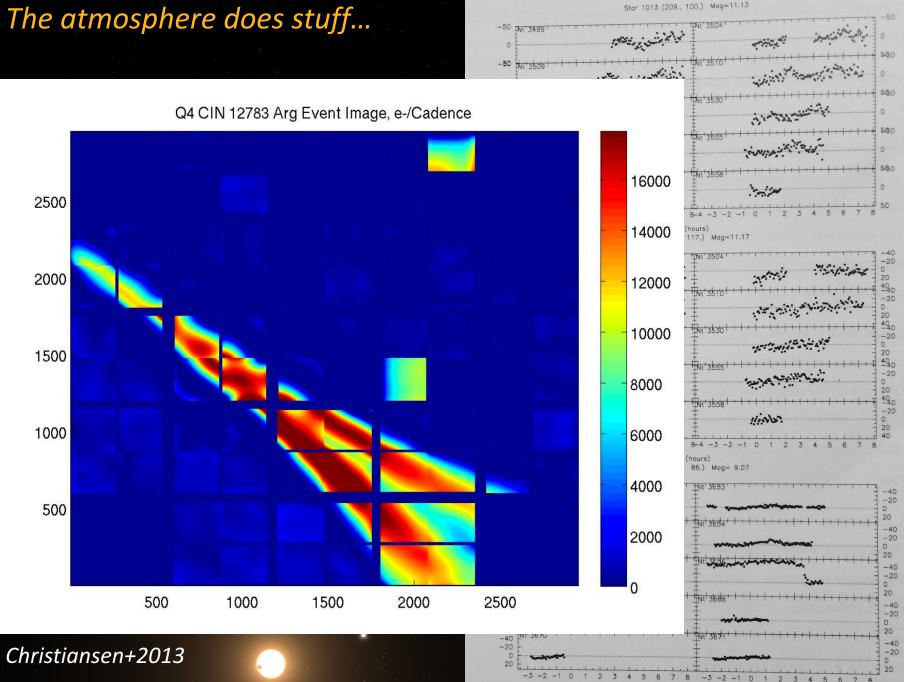
- Intra-pixel variations, e.g. Spitzer, K2 (not Kepler!)
- Inter-pixel variations, e.g. EPOCh





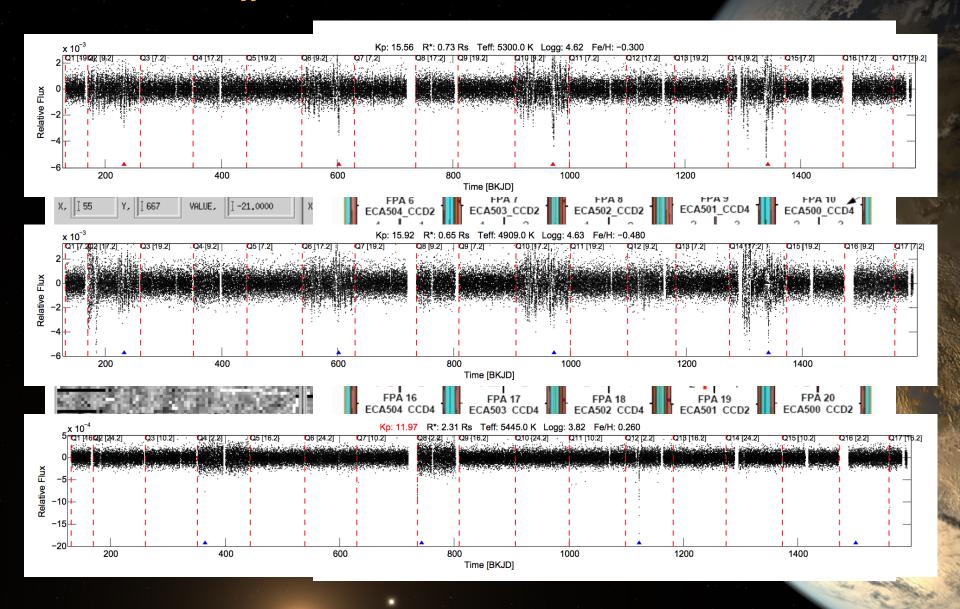
Cadence Index

The atmosphere does stuff...

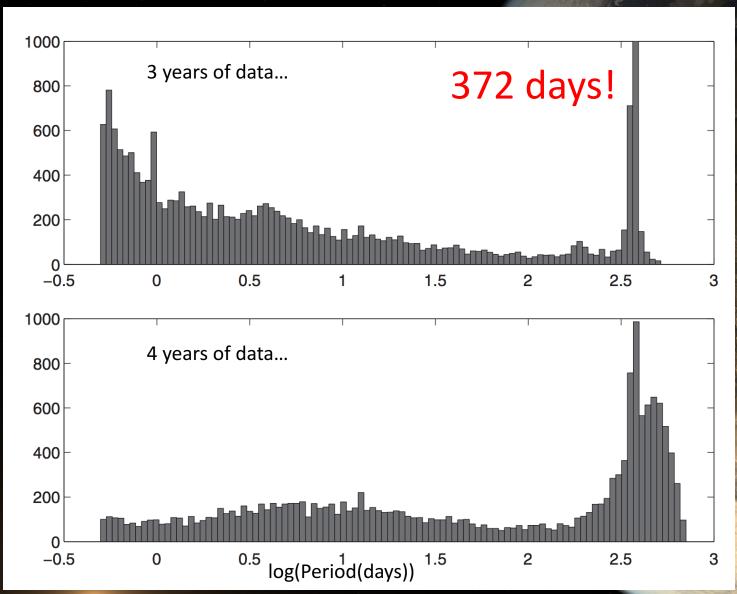


3.4 Aberrations
6.2.1 Description
6.2.2 Photometric Consequences
6.2.3 Mitigation
6.3 Video Crosstalk71
6.4 Science Clock and Video Crosstalk into FGS72
6.5 Start-of-Line Ringing72
6.5.1 Description
6.5.2 Photometric Consequences
6.5.3 Mitigation
6.6 Undershoot
6.6.1 Description
6.6.2 Photometric Consequences
6.6.3 Mitigation
6.7 Aliased High Frequency Noise
6.7.1 Aliased to near-Zero Spatial Frequency: Rolling Bands
6.7.2 Aliased to near-Nyquist Spatial Frequency: Moiré
6.7.3 Detection in Row Fourier Transforms
6.7.4 Excitation by Bright Sources
6.7.5 Photometric Consequences
6.7.6 Mitigation
6.8 Long Cadence Artifact Removal Pixels (ARPs)
5.2 Clocks and Sampling65

Detectors do stuff...



Detectors do stuff...



Summary

There is a lot standing between you and a clean transit signal!

You can make clever choices in order to minimise/mitigate/isolate noise sources (or at least the timescales of those noise sources)

- Target selection
- Stability of instrument
- The more the merrier, for treating common mode systematics)

For remaining (and sometimes unavoidable) noise sources, exoplaneteers have been relying on increasingly sophisticated noise models (e.g. Gaussian process) – see rest of #sagan2016!

