



InSb (3.6 & 4.5 μ m) Linearization

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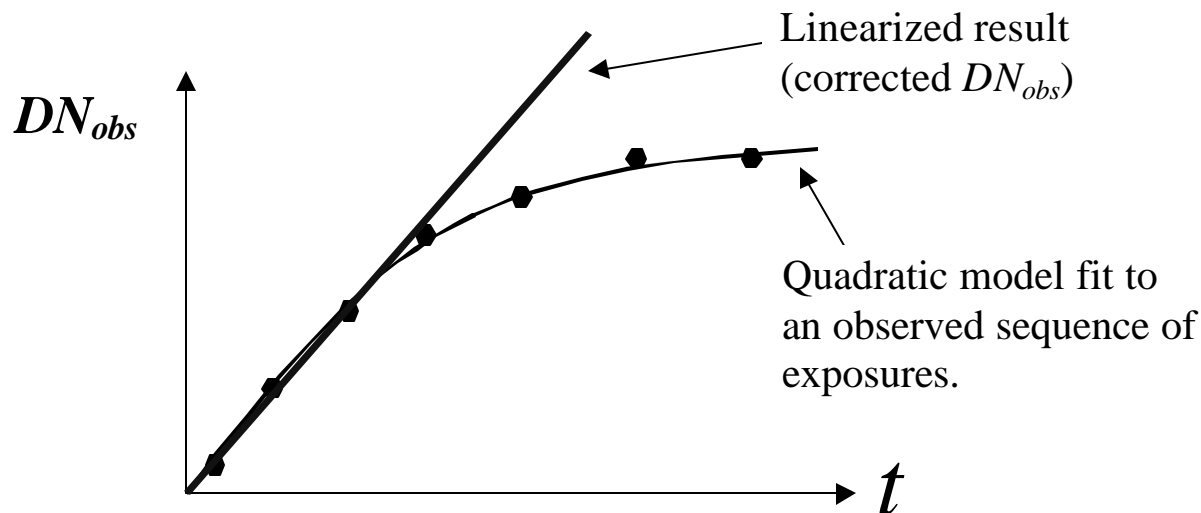
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- We have a new module (“Fowlinearize”) that will correct InSb (Channels 1 and 2) array data for non-linearity using a model that is a quadratic in time:

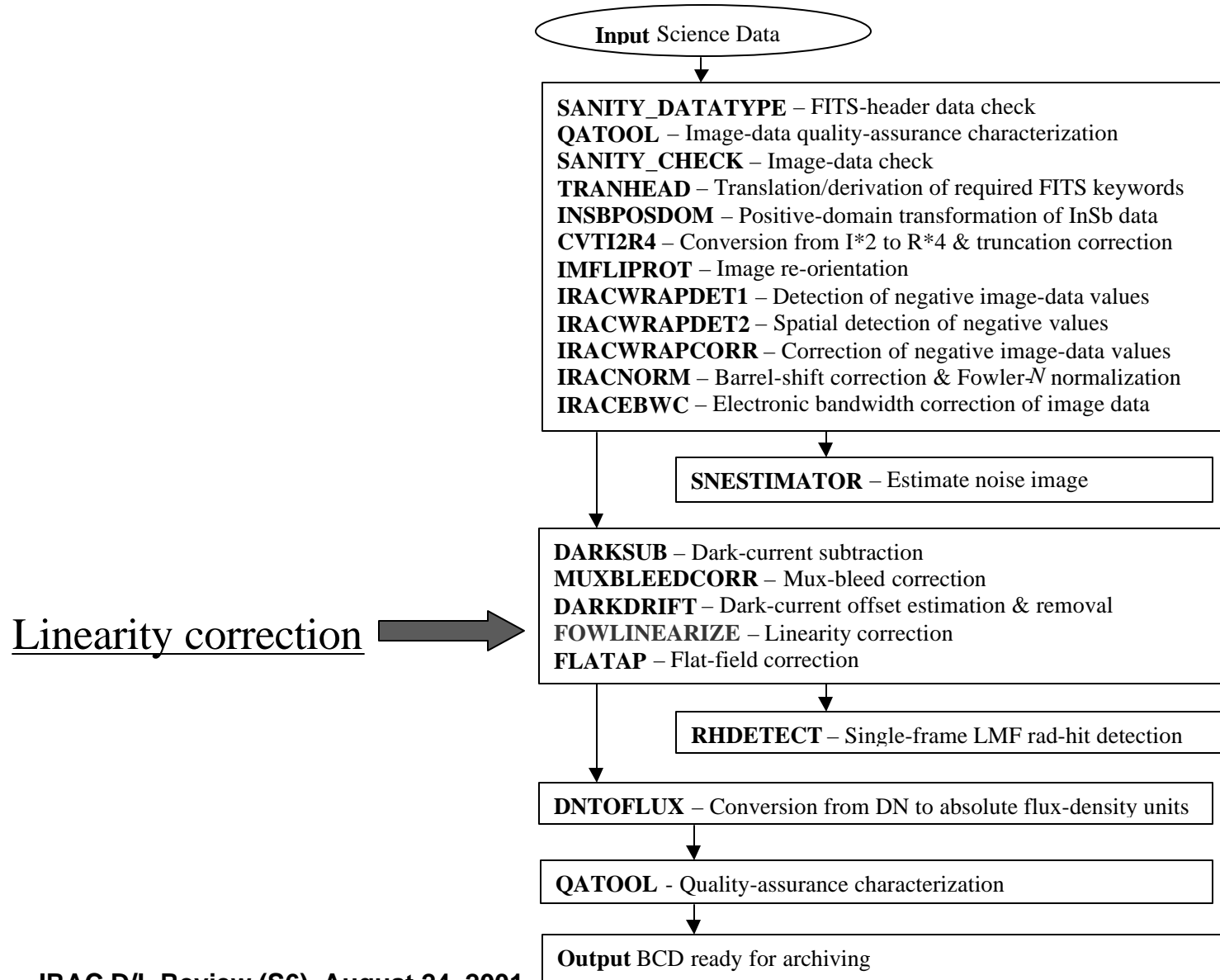
$$DN_{obs} = mt + At^2$$

- A quadratic function meets the 1% requirement over entire data range for the InSb flight array only.
- The coefficients of the above model are computed by the LINCAL module (in the Linearity Calibration Thread), where m = linearized count rate and A = quadratic coefficient.





Processing in Pipeline





- **INPUTS:**

- ◉ Single plane science DCE (FITS image).
- ◉ Non-linearity (quadratic) model coefficients for every pixel.
- ◉ Noise model image to use in error propagation.
- ◉ Mask images which specify: hot/dead pixels, saturated pixels, and pixels for which a non-linearity model could not be computed.

- **OUTPUTS:**

- ◉ New science DCE with linearized counts.
- ◉ Corresponding uncertainty image for the linearized DCE.
- ◉ Updated “dce-mask” image indicating which pixels could and could not be linearized.
- ◉ Output log file showing processing statistics and error messages.



- Inversion of the above quadratic includes a correction due to “Fowler sampling error”.
- Fowler sampling is used to reduce the effective read read noise. This causes the actual number of photons in the well to exceed that returned by the read-out electronics. Difference could be as large as 8%.
- The final solution for “true” linearized counts (DN) is as follows:

$$DN_{lin} = \frac{1 - \sqrt{1 - 4L_{cal}DN_{obs}}}{2L_{cal}}$$

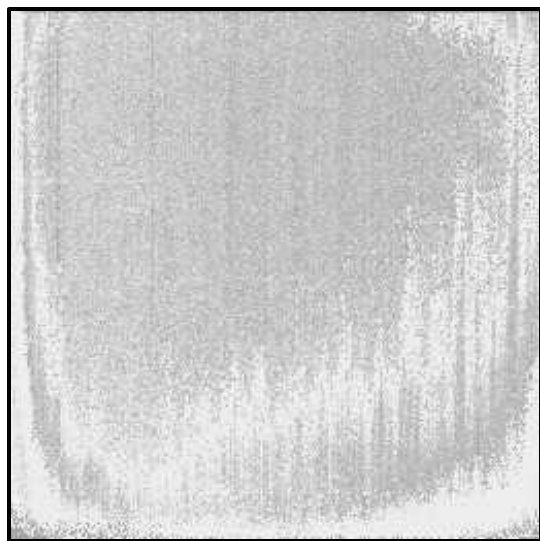
- Where the constant L_{cal} depends on the quadratic model coefficients (A, m), number of “Fowler sample” read-out times (n) and number of “wait” periods (w) between pedestal and signal read-out times:

$$L_{cal} = \left(\frac{A}{m^2} \right) \left(\frac{1}{n(w+n)^2} \right) \left[\sum_{i=1}^n i^2 - \sum_{i=w+n}^{w+2n} i^2 \right]$$

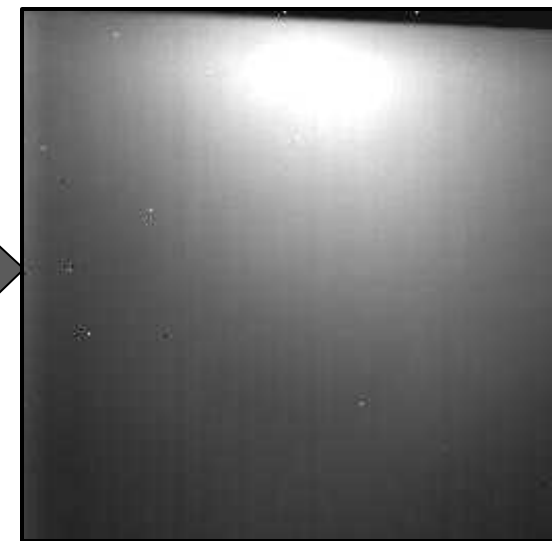
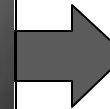
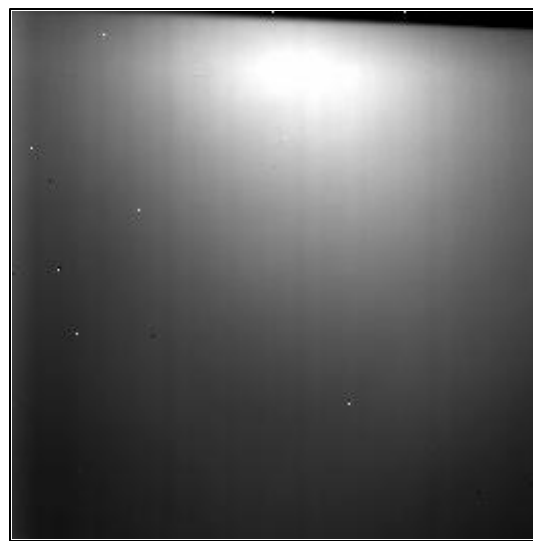


Image Input/Output Results

Testing on a transmission-calibrator flat in channel 1 (3.6 μ m) where all pixel counts are less than $\approx 1/4$ the saturation value.



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“Non-linearity model
image” of A-coefficients



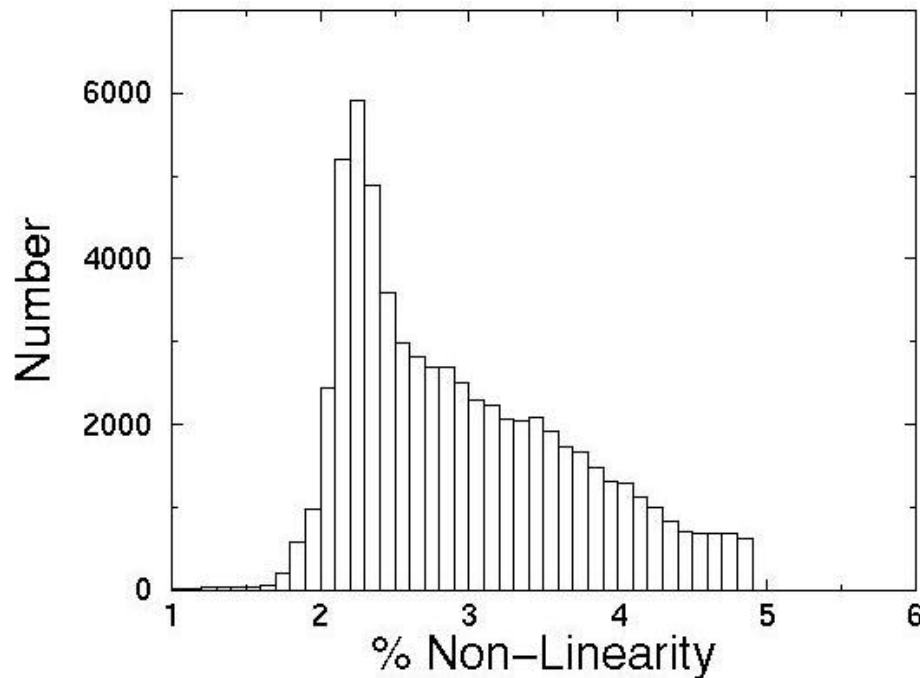
$$DN_{obs} = mt + At^2$$

Raw (un-linearized)

Linearized image

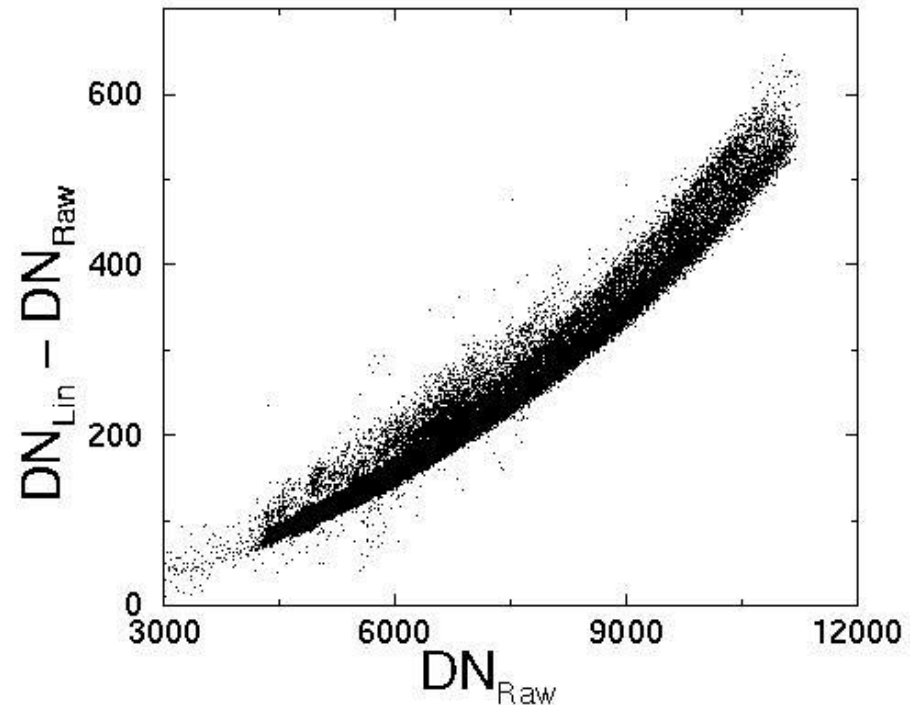


Distribution of pixel
non-linearities



$$\%NL = \frac{DN_{lin} - DN_{raw}}{DN_{raw}} \times 100$$

Linearized – Observed (raw)
counts vs. observed counts.





SiAs Array Linearization

- SiAs arrays show considerable non-linearity below 90% full well compared to the InSb arrays which are essentially linear.
- We require a cubic function to meet the 1% requirement for fitting SiAs linearity data (5.8 and 8 μm bands) over the entire dynamic range (see presentation by J. Surace):

$$DN_{obs} = At + Bt^2 + Ct^3$$

- This must be inverted to solve for t and the linearized counts are given by $DN_{lin} = At$. Of the three possible roots, instrument team has shown that the appropriate physical solution is that which satisfies constraints from characterization of the non-linearity.
- This code upgrade will be completed by end of S6.